

Odd Carbon Long Linear Chains $HC_{2n+1}H$ (n = 4-11): **Properties of the Neutrals and Radical Anions**

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Abstract: The optimized geometries, adiabatic electron affinities, vertical electron affinities, vertical electron detachment energies (for the anions), and IR-active vibrational frequencies have been predicted for the long linear carbon chains $HC_{2n+1}H$ (n = 4-11). The B3LYP density functional combined with DZP and TZ2P basis sets was used in this theoretical study. These methods have been extensively calibrated versus experiment for the prediction of electron affinities (Chem. Rev. 2002, 102, 231). The computed physical properties are discussed and compared with the even carbon chains HC₂₀H. The predicted electron affinities form a remarkably regular sequence: 2.12 eV (HC₉H), 2.42 eV (HC₁₁H), 2.66 eV (HC₁₃H), 2.85 eV (HC₁₅H), 3.01 eV (HC₁₇H), 3.14 eV (HC₁₉H), 3.25 eV (HC₂₁H), and 3.35 eV (HC₂₃H). These electron affinities are as much as 0.4 eV higher than those for analogous even carbon chains. The predicted structures display an intermediate cumulene-polyacetylene type of bonding, with the inner carbons appearing cumulenic and the outer carbons polyacetylenic.

I. Introduction

Increased attention has been given over the last two decades to the various carbon chains Cn, HCnH, CnH2, their isomers and ions, related nitrogen and sulfur containing molecules, and polyaromatic hydrocarbons.¹⁻⁷ These are still considered leading contenders among larger molecules possibly existing in interstellar space, but also as possible carriers of diffuse interstellar bands (DIBs) or unidentified interstellar infrared (UIR) emission bands.⁸⁻¹⁰ So far, the existence of 123 molecules¹¹ has been confirmed in interstellar space, but only a few of these have also been found in the DIBs' spectral region (400-900 nm). The largest molecule discovered in space, HC11N, is one of the reasons why the family of long carbon chains $HC_{2n+1}H$ is similarly expected to exist in the interstellar medium. Maier et al.¹² reported in 1998 that absorption bands of the carbon chain anion C_7^- match four of the DIBs exceedingly well. Despite critical comparisons of the laboratory data with new astrophysical observations^{13,14} the assignment of C_7^- remains controversial.

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Similarly controversial is the degree of contribution from long chain carbon molecules to both the DIB and UIR.

Another interesting question that has been raised concerns the size limit for the long carbon chains. Maier and co-workers^{15–17} succeeded in observing the neutral clusters C_{2n} (n =3-7) and C_{2n+1} (n = 2-10) and the anionic clusters C_{2n} (n = 2-10) 2–11) and C_{2n+1}^{-} (n = 2-10) using laser ablation of graphite with trapping in neon matrixes. Recently, electronic absorption spectra of neutral C_{2n+1} (n = 5-14) and anionic C_{2n} (n = 5-14) 5-18) in argon matrixes were presented by Szczepanski, Fuller, Ekern, and Vala,¹⁸ thus further increasing the range of the carbon chains covered in the quest for the possible DIB/UIR bands' carriers. In the case of polyacetylenic chains, electronic absorption spectra of $HC_{2n}H$ and $HC_{2n}H^-$ (n = 6-12) in neon matrices have been reported by Grutter, Wyss, Fulara, and Maier.¹⁹ More importantly, electronic spectra of the chains $HC_{2n}H$ (n = 8-13) have been recorded using the resonant two-color two-photon ionization spectroscopy in the gas phase,²⁰ the latter essential for the proper comparison of the laboratory and astronomical data. Even though the detection of polyacetylenic chains HC_nH by radio astronomy is hindered by the absence of permanent dipole moments, they are expected to be abundant in interstellar space.

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Our previous study²¹ focused on the description of the evenmembered structures $HC_{2n}H$ and $HC_{2n}H^-$ (n = 6-12). To complement this work, we turn our interest to the odd-membered structures $HC_{2n+1}H$ and $HC_{2n+1}H^-$ (n = 4-11). Fulara, Freivogel, Forney, and Maier²² pointed out in 1995 that the structures of these odd carbon species are linear and reported electronic transitions in the visible and near-infrared regions of the spectrum. Two years ago, the first gas-phase detection of linear HC11H and HC13H was provided by Ball, McCarthy, and Thaddeus using cavity ringdown spectroscopy.²³ Unfortunately, no satisfactory assignment to interstellar features has yet been made. To our knowledge, no experimental detection of long anionic carbon clusters $HC_{2n+1}H^{-}$ has been yet reported.

Following controversy due to discussions of the self-interaction problem,²⁴⁻³⁵ recent applications of density functional theory (DFT) methods to negatively charged molecules have shown several of these to be reliable methods for the prediction of adiabatic electron affinities. The self-interaction problem has been reputed to cause the DFT method to predict some anionic systems to be unbound relative to the analogous neutral systems.^{36,37} This effect has been tested, and it has been shown that even though the HOMO energies obtained from pure DFT functionals may have positive eigenvalues, DFT nevertheless makes reasonable predictions of the anion physical properties.³⁶ Rösch and Trickey speculated³⁷ that for large molecules in which the electron population is delocalized over the molecule, the self-interaction problem is delocalized and therefore less significant. With these factors in mind, optimized geometries, adiabatic electron affinities, vertical electron affinities, vertical electron detachment energies, and harmonic vibrational frequencies of $HC_{2n+1}H$ and $HC_{2n+1}H^-$ (n = 4-11) are presented in this work.

II. Theoretical Method

Total energies, equilibrium geometries, and harmonic vibrational frequencies were determined for the neutral and anionic species for each of the eight molecules studied. The B3LYP density functional, which is a generalized gradient approximation and employs the dynamical correlation functional of Lee, Yang, and Parr³⁸ in conjunction with Becke's three-parameter HF/DFT hybrid exchange functional (B3),³⁹ was used. We employed the double- ζ plus polarization (DZP) [9s5p2d/4s2p1d] and triple-ζ plus polarization (TZ2P) [10s6p2d/5s3p2d]

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contracted Gaussian basis set due to Dunning.⁴⁰ Accordingly, for HC₂₃H, the DZP and TZ2P basis sets contain 355 and 570 contracted Gaussian functions, respectively. The specific choice of the B3LYP density functional was based on extensive previous calibrative studies of electron affinities.⁴¹ It has been argued specifically that for the long carbon chains the B3LYP density functional should be used in order to predict the bond lengths properly.42,43

The quantum chemical computations of this study were conducted with the GAUSSIAN 9444 system of DFT programs. To maintain consistency, spin unrestricted Kohn-Sham orbitals were used for all computations. Both the neutral and anion geometries were optimized via analytic gradients of the B3LYP density functional. Numerical integration of the functional was carried out using the GAUSSIAN 94 default grid consisting of 75 radial shells with 302 angular points per shell. The mass-weighted force constant matrix and, hence, the harmonic vibrational frequencies were determined analytically.

III. Results and Discussion

One of the main goals of this study was to predict the fundamental properties of long odd-numbered carbon chains and compare them with the even-numbered chains. It has been shown before that properties of the smaller even- and oddnumbered structures can often differ extensively.45-48

In accordance with Hund's rules and experiment, we confirm that ${}^{3}\Sigma_{g}^{-}$ is the electronic ground state for all neutral molecules. This is due to the partially occupied HOMOs, (π_g^2) for HC_{4n+1}H (n = 2-5) and (π_{u}^{2}) for HC_{4n+3}H. An extra electron added to this molecular orbital leads to the ${}^{2}\Pi_{g}$ or ${}^{2}\Pi_{u}$ ground state of the anions $HC_{2n+1}H^{-}$ (n = 4-11). The electron configurations for the first few members of the series are as follows:

$${}^{3}\Sigma_{g}^{-} \text{HC}_{9}\text{H} \dots (10\sigma_{g}^{2})(9\sigma_{u}^{2})(1\pi_{u}^{4})(1\pi_{g}^{4})(2\pi_{u}^{4})(2\pi_{g}^{4})(3\pi_{u}^{2})$$

$${}^{2}\Pi_{u} \text{HC}_{9}\text{H}^{-} \dots (10\sigma_{g}^{2})(9\sigma_{u}^{2})(1\pi_{u}^{4})(1\pi_{g}^{4})(2\pi_{u}^{4})(2\pi_{g}^{4})(3\pi_{u}^{3})$$

$${}^{3}\Sigma_{g}^{-} \text{HC}_{11}\text{H} \dots (12\sigma_{g}^{2})(11\sigma_{u}^{2})(1\pi_{u}^{4})(1\pi_{g}^{4})(2\pi_{u}^{4})(2\pi_{g}^{4}) (3\pi_{g}^{2})$$

$$(3\pi_{u}^{4})(3\pi_{g}^{2})$$

$${}^{2}\Pi_{u} \text{HC}_{11}\text{H}^{-} \dots (12\sigma_{g}^{2})(11\sigma_{u}^{2})(1\pi_{u}^{4})(1\pi_{g}^{4})(2\pi_{u}^{4})(2\pi_{g}^{4})$$

$$\frac{11_{g} 11 c_{11} 11}{(3\pi_{u}^{4})(3\pi_{g}^{3})}$$

$${}^{3}\Sigma_{g}^{-} \text{HC}_{13}\text{H} \dots (14\sigma_{g}^{2})(13\sigma_{u}^{2})(1\pi_{u}^{4})(1\pi_{g}^{4})(2\pi_{u}^{4})(2\pi_{g}^{4})$$

$$(3\pi_{u}^{4})(3\pi_{g}^{4})(4\pi_{u}^{2})$$

$${}^{2}\Pi_{u} \operatorname{HC}_{13} \operatorname{H}^{-} \dots (14\sigma_{g}^{2})(13\sigma_{u}^{2})(1\pi_{u}^{4})(1\pi_{g}^{4})(2\pi_{u}^{4})(2\pi_{g}^{4}) (3\pi_{u}^{4})(3\pi_{g}^{4})(4\pi_{u}^{3})$$

The remainder of the pattern through HC23H and HC23H⁻ should be reasonably clear. We note that in the case of even carbon series HC_{2n}H the HOMO π -orbital is fully occupied.²¹

The optimized equilibrium geometries ($D_{\infty h}$ -symmetry) are all minima on their respective potential surfaces and are shown in Tables 1 and 2, where the C_1 in our numbering scheme begins with the carbon atom in the center of the chain structure. The numbering scheme is further documented by the optimized structure of HC₂₃H in Figure 1. The geometry optimizations in

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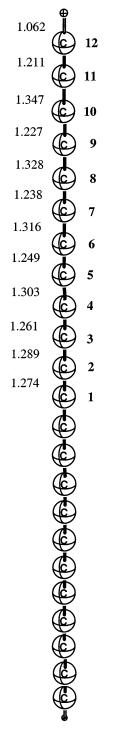


Figure 1. An example of atom numbering in chain structures, with TZ2P B3LYP optimized bond distances in Å for $HC_{23}H$.

the present study began with equivalent bond lengths between the carbon atoms, thus imposing a strictly cumulenic structure. Ultimately, we found significant bond length alternation in the fully optimized structures, which may be rationalized as a strong

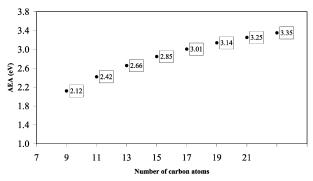


Figure 2. Odd carbon adiabatic electron affinities (AEA) vs the number of carbon atoms in the chain. The TZ2P B3LYP method was used.

competition between two resonance structures, cumulenic and polyacetylenic. Polyacetylenic type bond distance alternation begins to prevail as one moves outward along the chain. It is very interesting that the bond length alternation for the outermost carbons is equal to if not more pronounced than that in the evennumbered $HC_{2n}H$ (n = 6-12) species of our previous study.²¹

The fact that these chains have an odd number of carbons implies that, if one starts with triple/single bond alternation from the ends, there will occur a mismatch in the center of the chain. This mismatch is reminiscent of soliton formation in odd polyenes.^{49,50} In other words, the mismatch forces the switch from the acetylenic to the cumulenic structure. Note that the mismatch is of a different nature for the HC_(4n+1)H and HC_(4n+3)H chains, which explains the oscillation in the optimized central bonds, found in Tables 1 and 2, a feature not previously discussed. The fact that the less stable cumulenic structure is forced by the mismatch also likely explains the higher electron affinity of the odd carbon compounds with respect to the even structures.

It is commonly agreed that because of overlap and repulsion within the (π) HOMOs, the addition of hydrogen at the ends of the chain leads to changes in the bonding and structure.⁴⁶ In their work on the linear carbon cluster anions C_n^{-} (n = 2-10), Watts and Bartlett observed47 continuous shifts from cumulenic to acetylenic structures after the addition of an electron to the C_n chain. This trend was recently also confirmed in the extensive DFT (B3LYP/6-31G*) study for C_n and C_n^- (n = 2-20).⁴⁸ The addition of the second electron in the dianions C_n^{2-} increased the degree of bond length alternation substantially. We have, however, pointed out²¹ that the addition of an electron to the neutral HC_{2n}H (n = 6-12) structures, which are isoelectronic with C_n^{2-} , leads to a slight shift away from acetylenic character as the triple bonds lengthen and the single bond shortens with respect to those of neutral $HC_{2n}H$. In the present study, the addition of an electron appears to have a negligible effect on the bond length alternation when the structures of neutral molecules are compared with the anions $HC_{2n+1}H^-$. Rotational constants calculated from the equilibrium structures (TZ2P basis set) are also reported in Table 1.

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Table 1. Optimized Geometries (Bond Distances in Å) and Rotational Constants (MHz) for the ${}^{3}\Sigma_{g}{}^{-}$ Ground States of Neutral HC_{2n+1}H (B3LYP, DZP/TZ2P basis sets, n = 4-11)

	HC₀H	HC ₁₁ H	HC ₁₃ H	HC ₁₅ H	HC ₁₇ H	HC ₁₉ H	HC ₂₁ H	HC ₂₃ H
C ₁ C ₂	1.302/1.290	1.285/1.273	1.297/1.285	1.286/1.274	1.295/1.283	1.287/1.274	1.294/1.282	1.287/1.274
C_2C_3	1.268/1.254	1.314/1.302	1.275/1.261	1.306/1.295	1.278/1.265	1.302/1.291	1.280/1.267	1.300/1.289
C_3C_4	1.342/1.331	1.259/1.245	1.321/1.311	1.267/1.254	1.313/1.302	1.272/1.258	1.308/1.297	1.274/1.261
C_4C_5	1.236/1.220	1.347/1.336	1.253/1.239	1.327/1.317	1.262/1.248	1.318/1.308	1.267/1.253	1.313/1.303
C_5C_6		1.233/1.217	1.351/1.334	1.249/1.235	1.331/1.321	1.258/1.244	1.322/1.312	1.263/1.249
C_6C_7			1.231/1.215	1.353/1.342	1.246/1.232	1.334/1.324	1.255/1.241	1.326/1.316
C_7C_8				1.299/1.213	1.355/1.344	1.244/1.230	1.337/1.327	1.253/1.238
C_8C_9					1.228/1.213	1.356/1.346	1.243/1.22	1.338/1.329
C_9C_{10}						1.226/1.212	1.357/1.347	1.242/1.227
$C_{10}C_{11}$							1.227/1.212	1.358/1.347
$C_{11}C_{12}$								1.227/1.211
C(n+1)/2H	1.070/1.062	1.070/1.062	1.070/1.062	1.070/1.062	1.070/1.062	1.071/1.062	1.071/1.062	1.071/1.062
A	402.7	222.4	135.6	88.7	61.1	43.9	32.6	24.9

Table 2. Optimized Geometries (Bond Distances in Å) for the ${}^{2}\Pi$ Ground States of the Anions HC_{2n+1}H⁻ (B3LYP, DZP/TZ2P basis sets, n = 4-11)

	HC ₉ H ⁻	HC ₁₁ H [−]	HC ₁₃ H ⁻	HC ₁₅ H ⁻	HC ₁₇ H ⁻	HC ₁₉ H ⁻	HC ₂₁ H ⁻	HC ₂₃ H-
C_1C_2	1.305/1.293	1.289/1.277	1.300/1.288	1.289/1.277	1.297/1.285	1.289/1.276	1.295/1.283	1.289/1.276
C_2C_3	1.274/1.261	1.315/1.304	1.279/1.266	1.307/1.296	1.281/1.268	1.303/1.291	1.282/1.270	1.301/1.289
C_3C_4	1.343/1.331	1.265/1.251	1.322/1.311	1.271/1.258	1.313/1.302	1.250/1.262	1.308/1.297	1.277/1.264
C_4C_5	1.244/1.229	1.345/1.336	1.258/1.245	1.327/1.316	1.266/1.253	1.318/1.307	1.270/1.257	1.312/1.302
C_5C_6		1.239/1.224	1.351/1.340	1.254/1.241	1.331/1.320	1.262/1.249	1.322/1.311	1.266/1.253
C_6C_7			1.236/1.221	1.353/1.342	1.251/1.237	1.334/1.323	1.259/1.245	1.325/1.314
C_7C_8				1.234/1.219	1.355/1.344	1.248/1.235	1.336/1.326	1.256/1.243
C_8C_9					1.232/1.218	1.356/1.345	1.247/1.233	1.338/1.327
C_9C_{10}						1.232/1.216	1.357/1.346	1.245/1.231
$C_{10}C_{11}$							1.230/1.215	1.358/1.347
$C_{11}C_{12}$								1.223/1.215
$C_{(n+1)/2}H$	1.067/1.059	1.067/1.059	1.067/1.059	1.068/1.059	1.068/1.060	1.068/1.060	1.068/1.060	1.069/1.060

Table 3. Theoretically Predicted IR-Active Harmonic Vibrational Frequencies (cm⁻¹) and Intensities (km/mol) for the X ${}^{3}\Sigma_{g}^{-}$ Ground States of HC_{2n+1}H (n = 4-11)^{*a*}

vibrational								
mode	HC₀H	HC ₁₁ H	HC ₁₃ H	HC ₁₅ H	HC ₁₇ H	HC ₁₉ H	HC ₂₁ H	HC ₂₃ H
$\pi_{ m u}$	223 (3)	166 (3)	124 (2)	96 (~2)	76 (~1)	61 (~1)	50 (~1)	42 (~1)
$\pi_{ m u}$	44 (~2)	37 (~2)	26 (~2)	216 (2)	17 (~2)	14 (~2)	12 (~1)	10 (~1)
$\pi_{ m u}$		47 (~1)	42 (~1)	37 (~1)	28 (~1)	24 (~2)	20 (~2)	17 (~2)
$\pi_{ m u}$				45 (~1)	41 (~1)	36 (~1)	29 (~1)	26 (~2)
$\pi_{ m u}$					47 (~1)	44 (~1)	40 (~1)	43 (~1)
$\pi_{ m u}$							46 (~1)	48 (~1)
$\pi_{ m u}$	56 (96)	59 (95)	64 (94)	61 (93)	62 (93)	62 (92)	62 (92)	63 (92)
$\sigma_{\rm u}$	85 (11)	71 (14)	60 (16)	53 (18)	47 (19)	42 (20)	38 (21)	34 (22)
σ_{u}	1536 (12)	13 (~2)	116 (5)	102 (8)	918 (9)	82 (10)	75 (11)	69 (11)
$\sigma_{ m u}$	1735 (6)	1440 (25)	1342 (34)	1251 (45)	1166 (55)	1087 (66)	1013 (76)	944 (86)
$\sigma_{ m u}$	2085 (9)	2028 (10)	1779 (3)	1468 (4)	1323 (7)	1204 (9)	1102 (11)	1016 (12)
$\sigma_{ m u}$	3462 (409)	2123 (6)	2092 (16)	2022 (7)	1797 (~2)	1530 (2)	1413 (5)	1312 (6)
$\sigma_{ m u}$		3463 (506)	2134 (4)	2118 (31)	2069 (6)	2016 (5)	1805 (~1)	1586 (~2)
$\sigma_{ m u}$			3464 (602)	2144 (13)	2110 (15)	2062 (~1)	2036 (4)	2004 (2)
$\sigma_{ m u}$				3465 (696)	2173 (10)	2128 (27)	2072 (17)	2012 (15)
σ_{u}					3465 (788)	2191 (7)	2153 (23)	2103 (21)
$\sigma_{ m u}$						3466 (877)	2201 (6)	2168 (19)
$\sigma_{ m u}$							3466 (~1000)	2208 (6)
$\sigma_{\rm u}$								3466 (~1000)

^a The TZ2P B3LYP method was used.

Theoretically predicted IR-active harmonic vibrational frequencies together with infrared intensities for the TZ2P basis set are presented in Tables 3 and 4. No scale factors were used to correct the predicted vibrational frequencies. These should serve as a guide for identifying the basic features in infrared spectra that might be attributed to the molecules under study. The strongest modes for the neutral species are predicted to be the highest (π_u) bending vibration and antisymmetric CH stretch (σ_u). While the CH stretch exhibits a constant frequency (3464 ± 2 cm⁻¹ for the neutrals; 3481 ± 3 cm⁻¹ for the anions) for all molecules, the wavelength of the (π_u) bending vibration features a continuous shift from 568 cm⁻¹ for HC₉H up to 632 cm⁻¹ for HC₂₃H. This is a very real trend, confirmed for smaller systems by earlier results of Seburg, McMahon, Stanton, and Gauss,⁵¹ in which the corresponding mode for the HC₅H molecule predicted at the CCSD/DZP level was 440 cm⁻¹. Fulara et al.²² reported earlier electronic absorption spectra for linear carbon chains HC_{2n+1}H (n = 2-7) in neon matrixes with characteristic vibronic progressions in the CC stretching mode (1961–1973 cm⁻¹). However, an assignment of several of the

⁽⁵¹⁾ Seburg, R. A.; McMahon, R. J.; Stanton, J. F.; Gauss, J. J. Am. Chem. Soc. **1997**, 119, 10838.

Table 4. Theoretically Predicted IR-Active Harmonic Vibrational Frequencies (cm⁻¹) and Intensities (km/mol) for the X $^{2}\Pi$ Ground States of HC_{2n+1}H⁻ (n = 4-11)^a

vibrational								
mode	HC₀H [−]	HC ₁₁ H ⁻	HC ₁₃ H ⁻	HC ₁₅ H ⁻	HC ₁₇ H ⁻	HC ₁₉ H ⁻	HC ₂₁ H ⁻	HC ₂₃ H ⁻
$\pi_{ m u}$	115 (200)	151 (~2)	118 (~1)	93 (~1)	74 (~1)	60 (~1)	119 (~1)	99 (~1)
$\pi_{ m u}$	144 (9)	294 (181)	210 (~1)	205 (~1)	170 (~1)	141 (~1)	197 (~1)	176 (~1)
$\pi_{ m u}$	233 (~1)	420 (~2)	272 (~1)	424 (40)	292 (~1)	236 (~1)	212 (~1)	245 (~1)
$\pi_{ m u}$	445 (12)	455 (13)	379 (155)	430 (97)	462 (97)	475 (7)	483 (~2)	268 (~1)
$\pi_{ m u}$	508 (116)	531 (91)	445 (5)	457 (9)	471 (5)	485 (4)	491 (~2)	489 (2)
$\pi_{ m u}$	536 (3)	541 (26)	466 (14)	473 (16)	481 (30)	491 (110)	504 (62)	496 (3)
$\pi_{ m u}$		711 (14)	540 (17)	494 (~1)	503 (6)	511 (26)	524 (73)	512 (29)
$\pi_{ m u}$					543 (3)	544 (2)	546 (2)	536 (99)
$\pi_{ m u}$			553 (95)	563 (105)	572 (105)	580 (105)	585 (104)	590 (104)
$\sigma_{ m u}$	855 (19)	1349 (59)	607 (11)	1027 (52)	470 (5)	423 (2)	754 (61)	692 (67)
$\sigma_{ m u}$	1706 (>1000)	1680 (>1000)	1166 (53)	1469 (53)	917 (53)	827 (56)	1102 (66)	1015 (62)
$\sigma_{ m u}$	1763 (>1000)	2009 (51)	1606 (>1000)	1536 (>1000)	1323 (89)	1203 (72)	1344 (>1000)	1286 (>1000)
σ_{u}	2078 (389)	2129 (711)	2079 (205)	2008 (123)	1469 (>1000)	1404 (>1000)	1415 (22)	1312 (458)
σ_{u}	3484 (450)	3483 (600)	2148 (~1000)	2129 (142)	1794 (4)	2005 (102)	1804 (4)	1574 (2)
σ_{u}			3482 (~1000)	2148 (>1000)	2065 (504)	2094 (>1000)	2057 (>1000)	1999 (113)
$\sigma_{\rm u}$				3482 (~1000)	2135 (>1000)	2128 (535)	2086 (>1000)	2058 (>1000)
$\sigma_{ m u}$					2167 (772)	2189 (756)	2152 (2)	2174 (17)
σ_{u}					3481 (>1000)	3479 (>1000)	2199 (862)	2204 (~1000)
σ_{u}							3479 (>1000)	3478 (>1000)

^a The TZ2P B3LYP method was used.

Table 5. Adiabatic Electron Affinities (AEA), Vertical Electron Affinities (VEA), and Anion Vertical Detachment Energies (VDE) in eV for HC_{2n+1}H [${}^{3}\Sigma_{g}^{-}$], HC_{2n+1}H⁻ [${}^{2}\Pi$] (n = 4-11)^a

		3			,	
molecule	AEA (DZP)	AEA (TZ2P)	VEA (DZP)	VEA (TZ2P)	VDE (DZP)	VDE (TZ2P)
HC ₉ H	2.03	2.12	2.02	2.11	2.04	2.13
$HC_{11}H$	2.35	2.42	2.34	2.41	2.36	2.42
HC13H	2.60	2.66	2.59	2.65	2.60	2.67
HC15H	2.80	2.85	2.79	2.84	2.80	2.85
HC17H	2.96	3.01	2.95	3.00	2.96	3.01
HC ₁₉ H	3.10	3.14	3.09	3.13	3.10	3.14
HC21H	3.21	3.25	3.21	3.24	3.22	3.29
$HC_{23}H$	3.31	3.35	3.31	3.34	3.32	3.37

^{*a*} The B3LYP method was used throughout; adiabatic EAs are not corrected for zero-point vibrational energies.

modes was not made. On the basis of the present results, we speculate that the vibration denoted $\nu_{(a)}$ at 619 cm⁻¹ for HC₁₃H²² can be assigned to the C-H bending mode. In addition to the two above-discussed strong IR-active harmonic vibrational frequencies, a number of weaker CC stretching modes (σ_u symmetry) and bending modes is listed. The vibronic band, which corresponds to the excitation of the triple bond stretch, has been reported by Ball, McCarthy, and Thaddeus²³ to lie at about 1950 cm⁻¹ for HC₇H, HC₉H, and HC₁₁H. Harmonic vibrational frequencies obtained for the anions $HC_{2n+1}H^-$ (Table 4) suggest that the infrared spectrum should be dominated mostly by CH stretching over 3000 cm⁻¹, as in a case of the neutral molecules, and by a number of strong CC stretching modes above 2000 cm⁻¹. The last strong CC stretching mode predicted here progresses from $\sim 1700 \text{ cm}^{-1}$ for HC₉H⁻ to $\sim 1300 \text{ cm}^{-1}$ for HC₂₃H⁻.

Table 5 summarizes the predicted adiabatic electron affinities (AEAs) and vertical electron affinities (VEAs) for all molecules studied, and the vertical electron detachment energies (VDEs) for the corresponding anions. The trend of increasing AEA as

the number of carbon atoms in the chain increases is further documented in Figure 2. The relatively high odd carbon AEAs exceed those predicted for the even-numbered HC_{2n}H series by ~0.4 eV (the value obtained²¹ for HC₂₄H was 2.95 eV). This supports the hypothesis that even HC_{2n+1}H⁻ can possibly exist in the interstellar medium. Finally, for comparison, we note that, for the linear C_n family, even-numbered structures have been reported to have higher AEAs than odd-numbered ones.^{47,48} Thus, the trend with alternating carbons for HC_nH is the opposite of that for C_n.

IV. Conclusions

For both neutral and anion radical species $HC_{2n+1}H$ and $HC_{2n+1}H^{-}$ (n = 4-11), we observe intermediate cumulenepolyacetylene type bonding with pronounced bond length alternation as the length of the carbon chain increases. The shorter carbon-carbon bonds at the chain ends take on triple bond character as their length decreases from ~ 1.29 to ~ 1.21 Å. We demonstrate that the electronic ground states are ${}^{3}\Sigma_{g}^{-}$ for the neutral molecules and ${}^{2}\Pi$ for the radical anions. Substantially higher adiabatic electron affinities, when compared with the even-numbered carbon chains HC2nH, are another example of the differences between odd- and even-membered structures. These conclusions support the hypothesis that HC_nH⁻ (even or odd carbon) may exist in interstellar media. The IRactive vibrational frequencies and the patterns of the IR spectra, presented for the first time for these molecules, may help to identify $HC_{2n+1}H$ and $HC_{2n+1}H^-$ in the laboratory in future gasphase experiments and possibly lead to their confirmation or rejection as DIB carriers.

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